

A composite image featuring three celestial bodies: Earth on the left, the Moon in the center, and Mars on the right. The InSight Mars rover is shown on the surface of Mars. The Earth shows the Americas, the Moon shows its craters, and Mars shows its reddish terrain and polar ice caps.

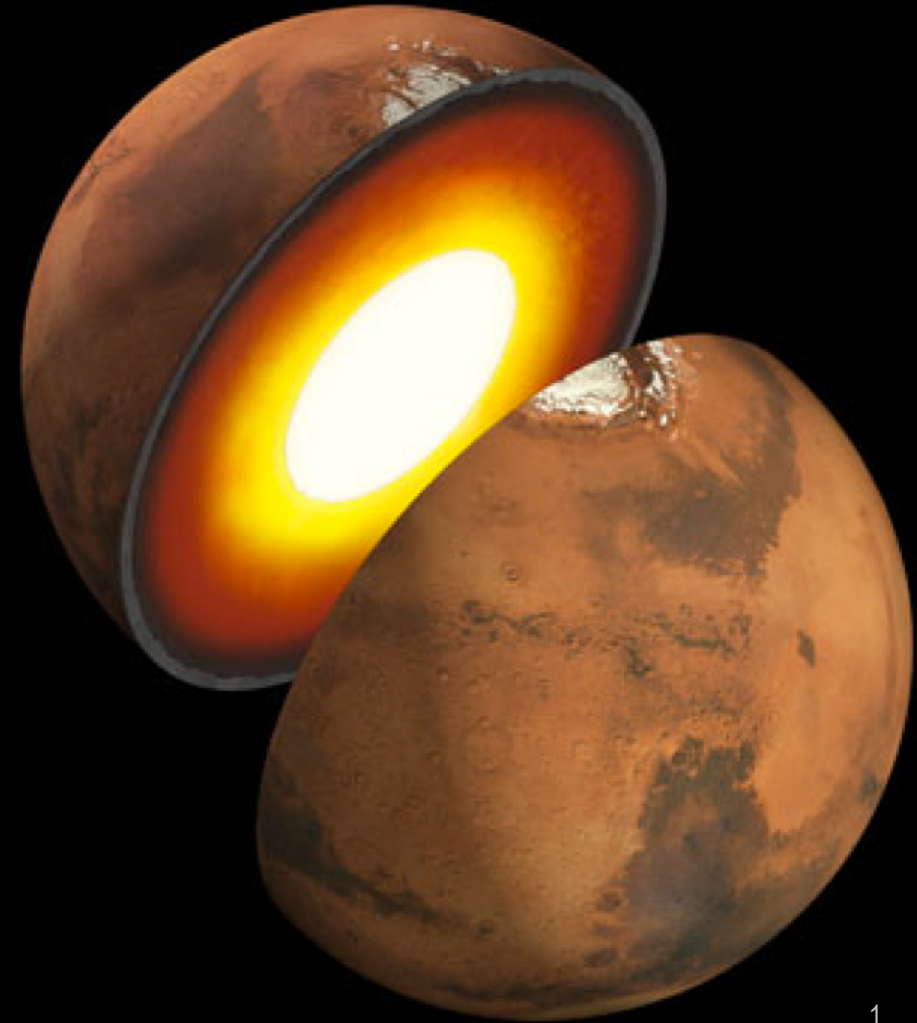
InSight Mission Science

W. Bruce Banerdt
Jet Propulsion Laboratory/California Institute of Technology
2 March, 2020

Goal: Provide constraints on the formation and early evolution processes of terrestrial planets by studying the internal structure of Mars.

In order to address this goal, InSight will determine, through geophysical measurements:

- ☐ Crustal thickness and large-scale layering
- ☐ Mantle structure
- ☐ Core size and density
- ☐ Global heat flux
- ☐ Rate and distribution of seismic activity
- ☐ Rate of meteorite impacts



- InSight travels back more than a hundred years, to Earth science at the dawn of the 20th century, to answer basic questions about the planet:
 - What is the thickness of the crust?
 - What is the structure of the mantle?
 - What is the size and density of the core?
 - What is the distribution of seismicity?
 - What is the heat flow from the planet?

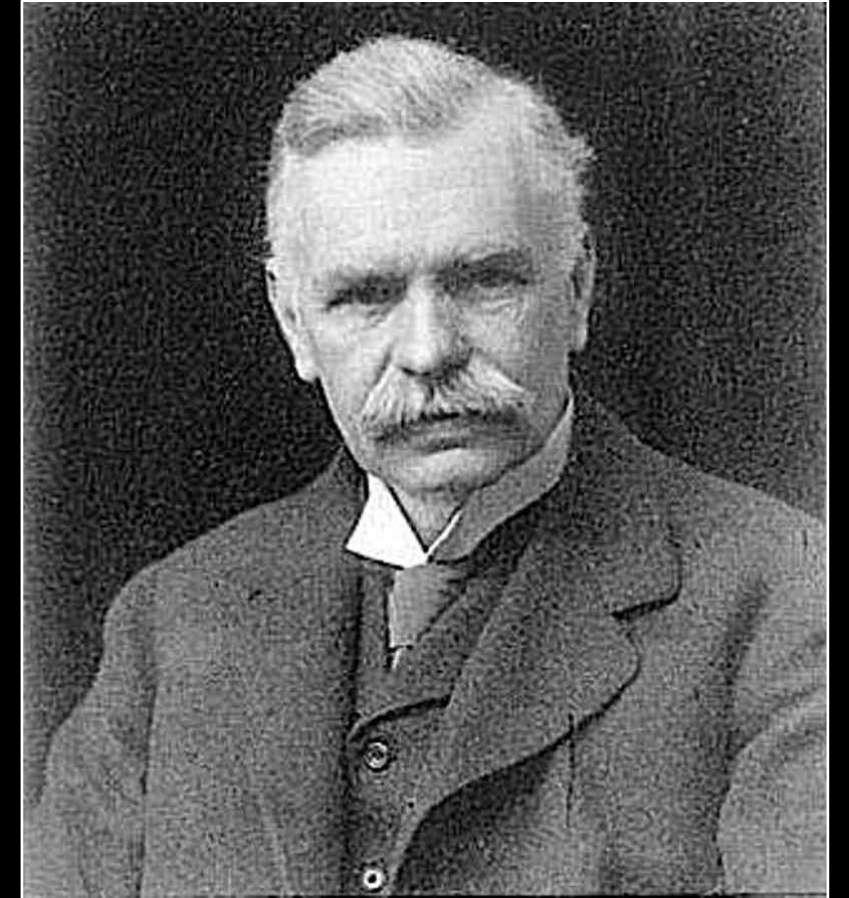
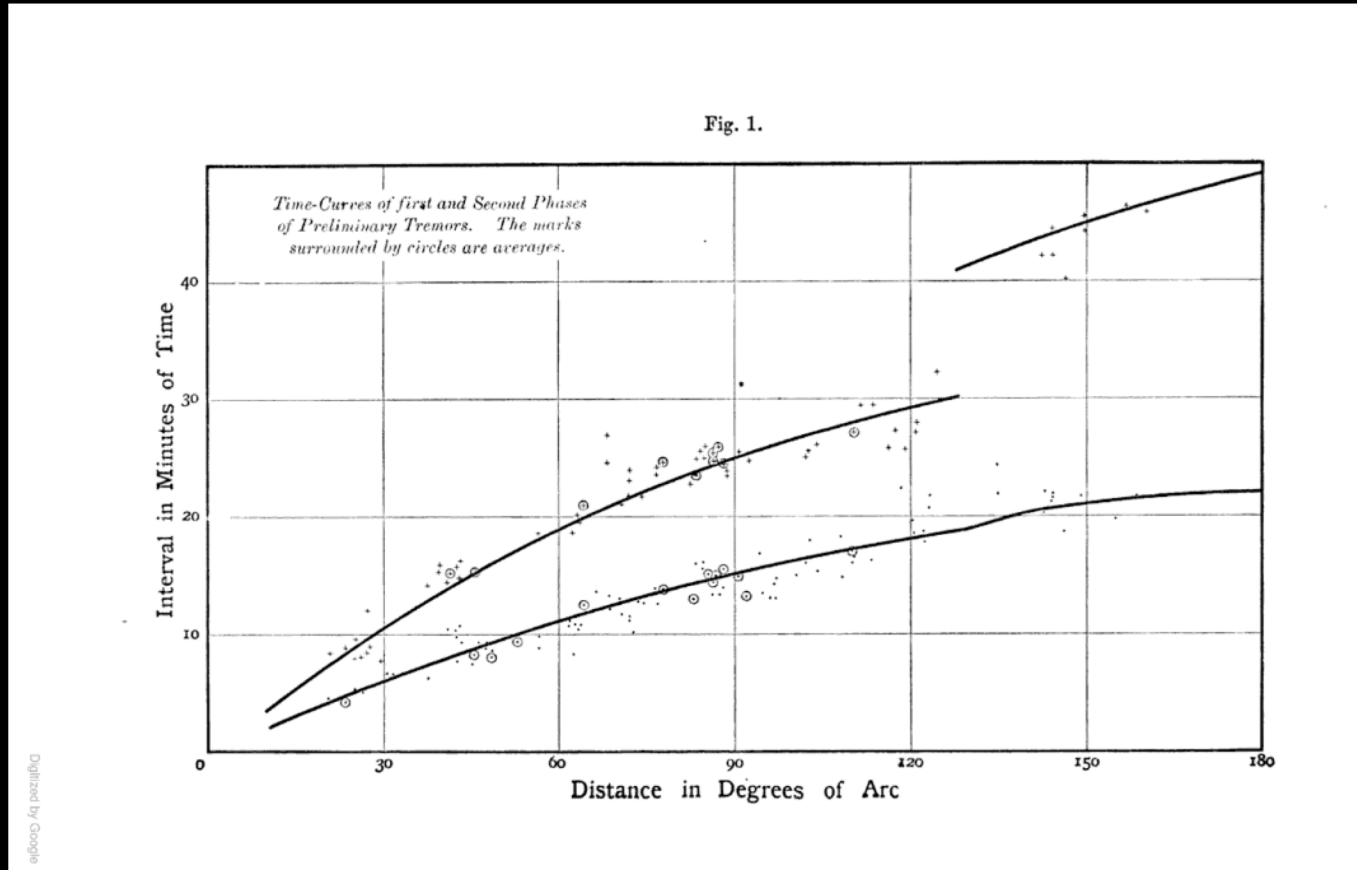


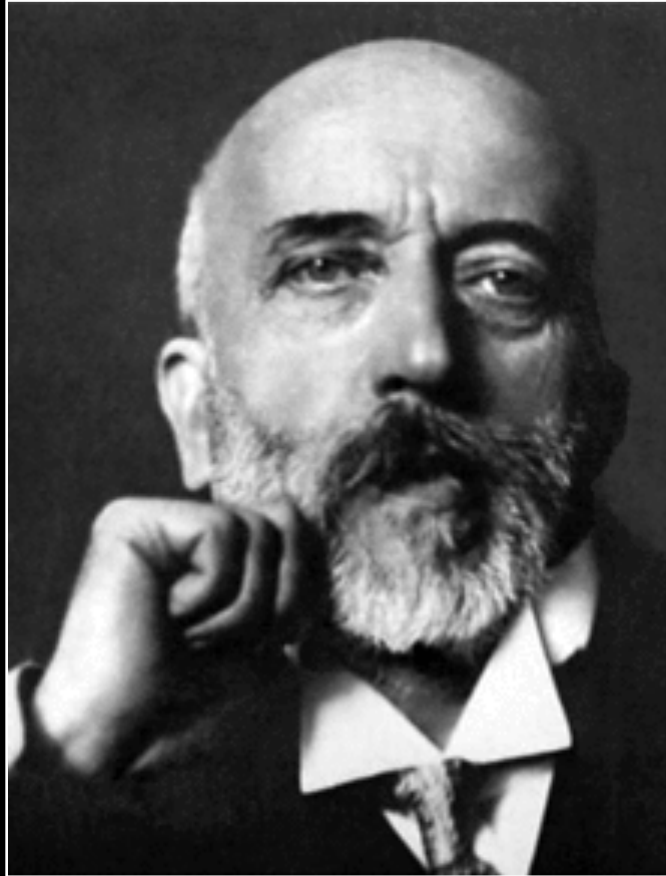
1889 – First Remote Detection of an Earthquake

- Ernst von Rebeür-Paschwitz
- A Tokyo earthquake was measured in Potsdam – 8850 km away.



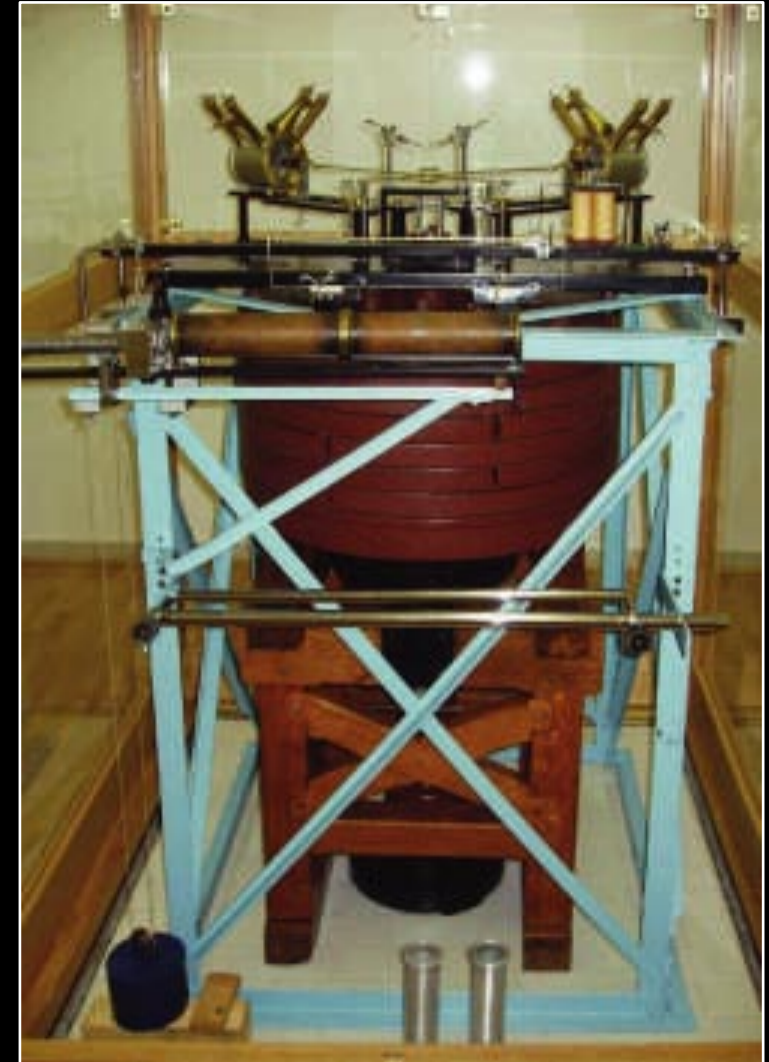
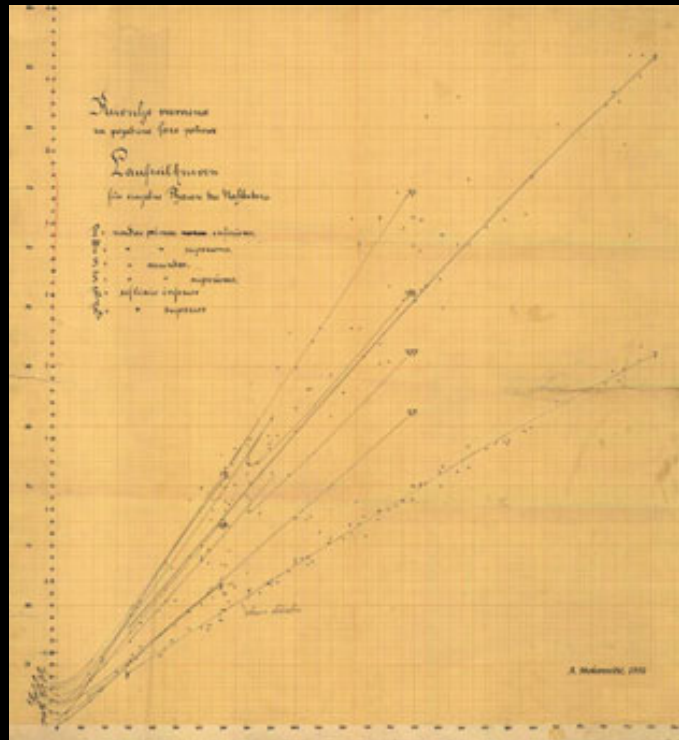
- Richard Dixon Oldham
- Derived from the interpretation of earthquake travel times.



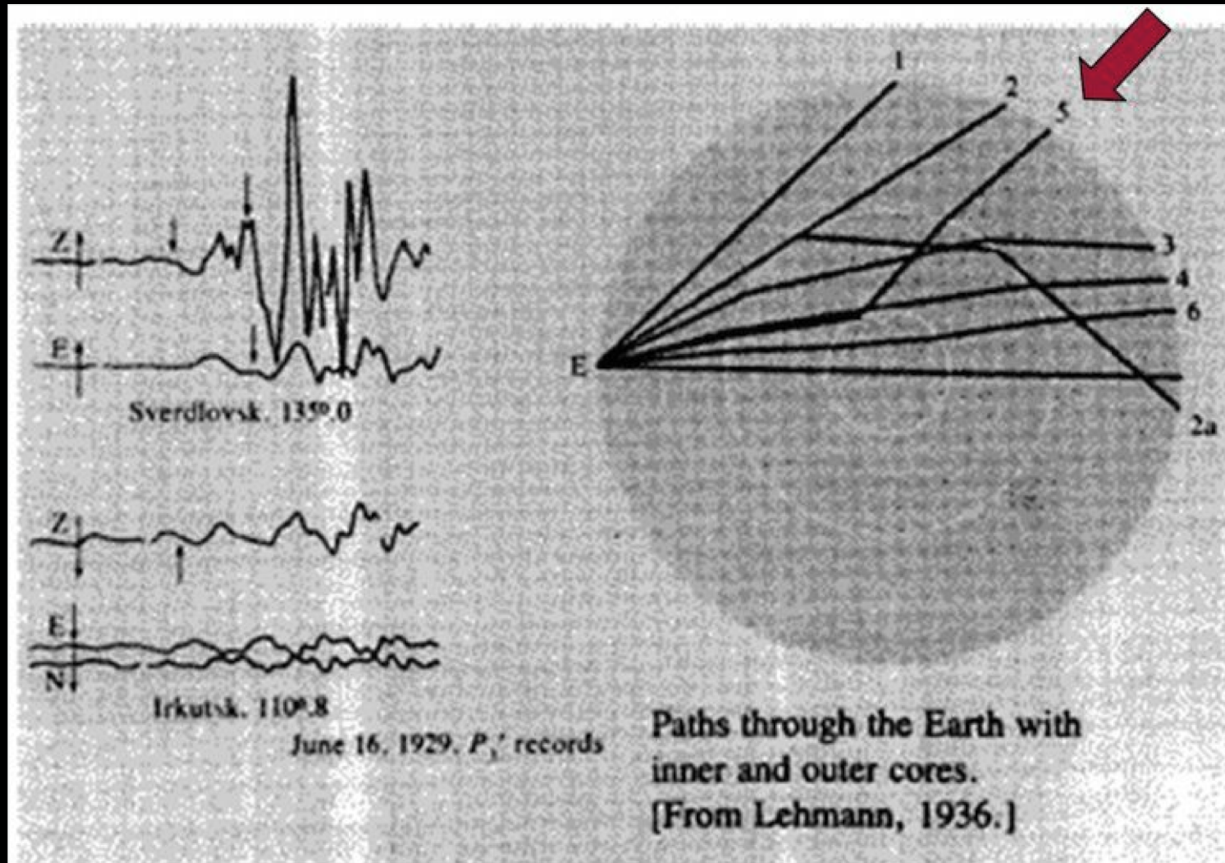


Andrija Mohorovičić

- Kupa Valley quake of 8 October 1909
- 50 km thickness beneath Croatia



Wiechert horizontal seismograph



Inge Lehman



Gutenberg and Richter, 1941
“Seismicity of the Earth”

In 50 years, seismologists went from the first detection of an earthquake to the basic understanding of the planet Earth.

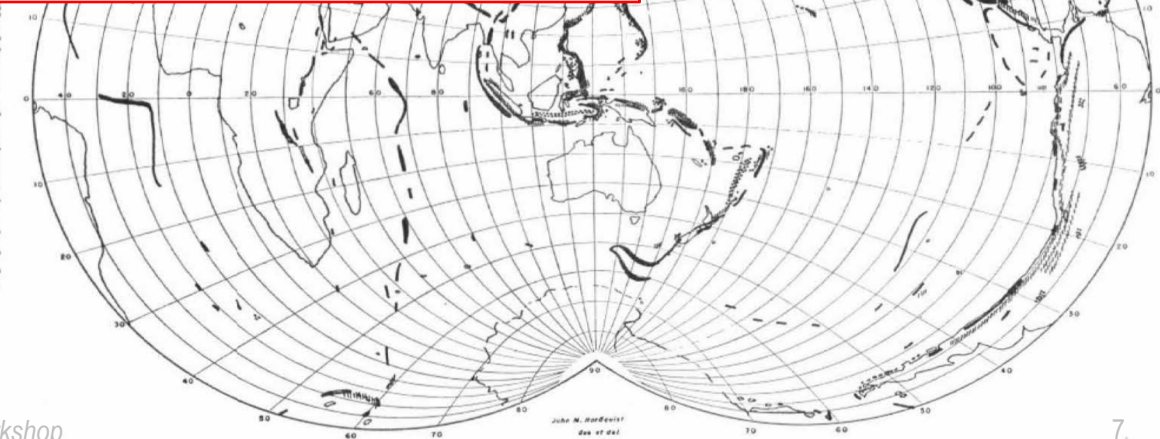
InSight plans to do this for Mars in 2 years!



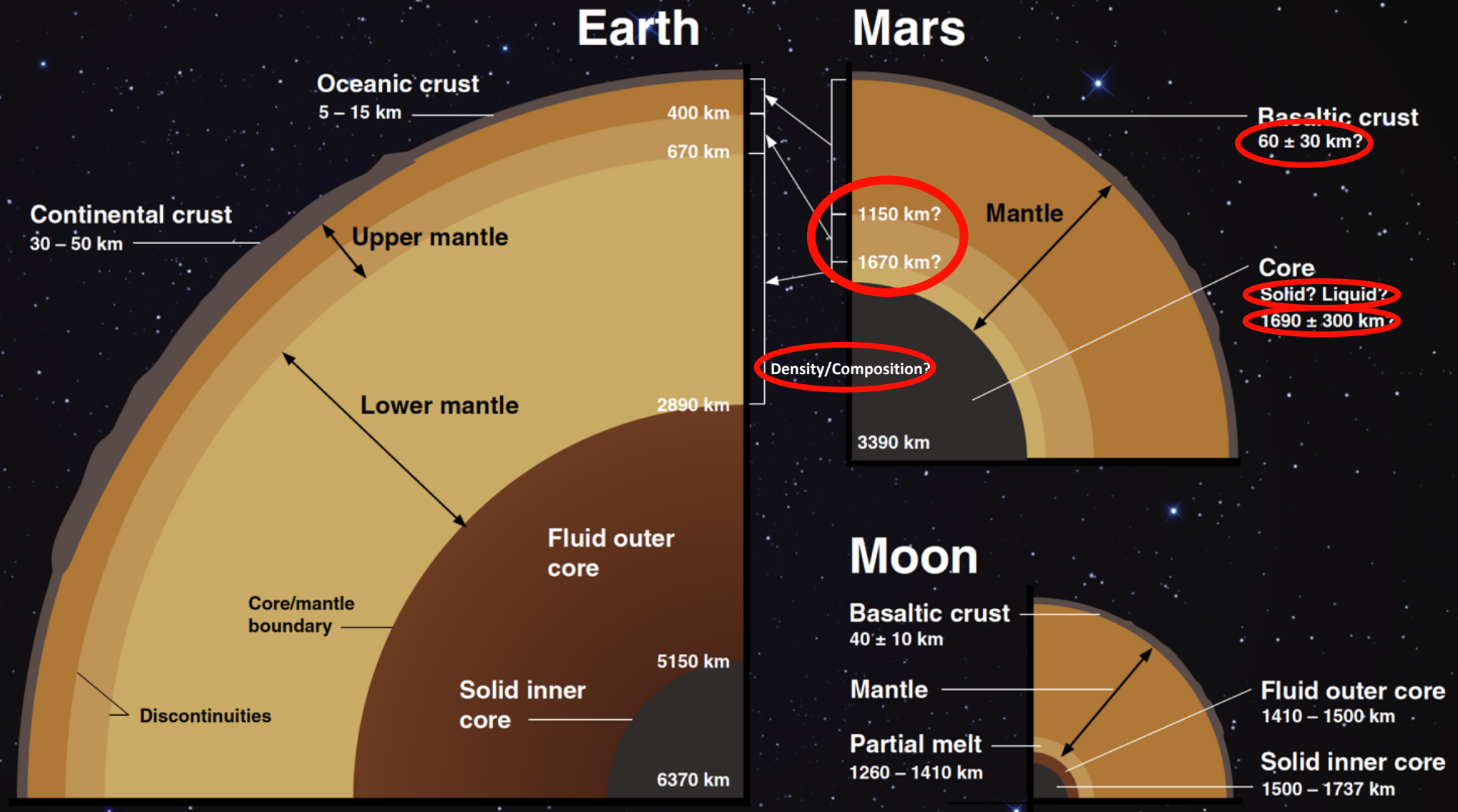
Wood-Anderson seismometer

2 March 2020

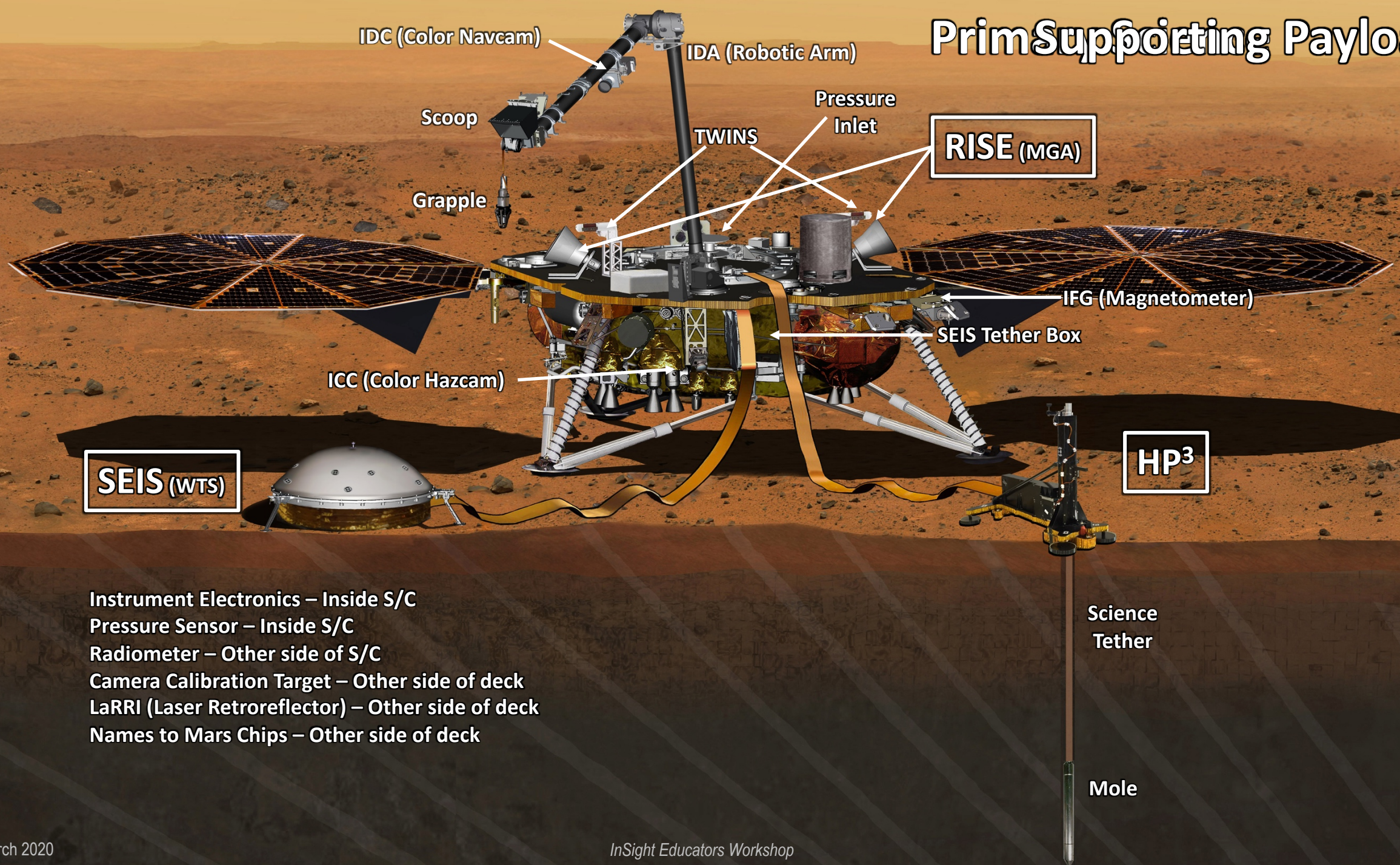
World map showing seismic belts



Mars Structure Compared to Earth and Moon



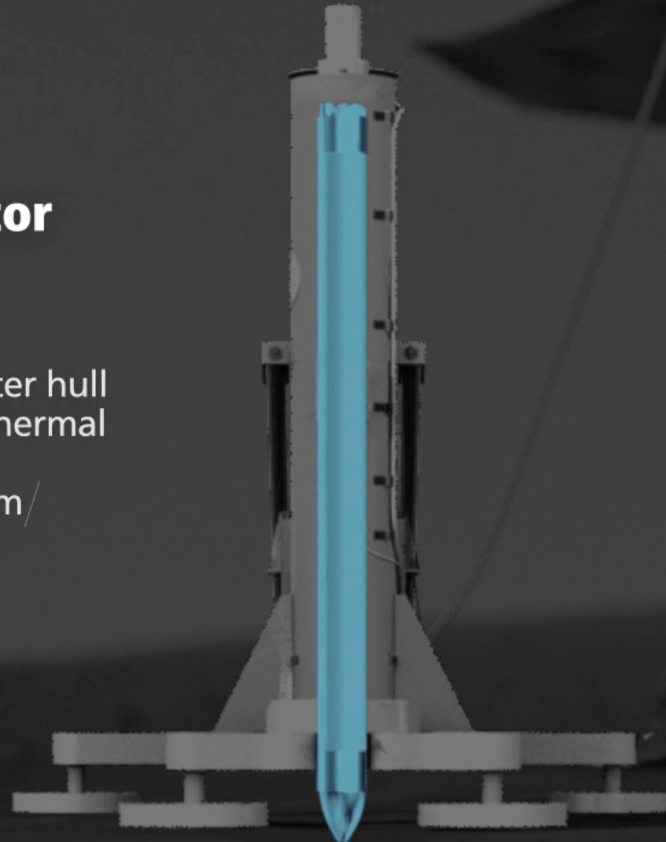
Primary Supporting Payload

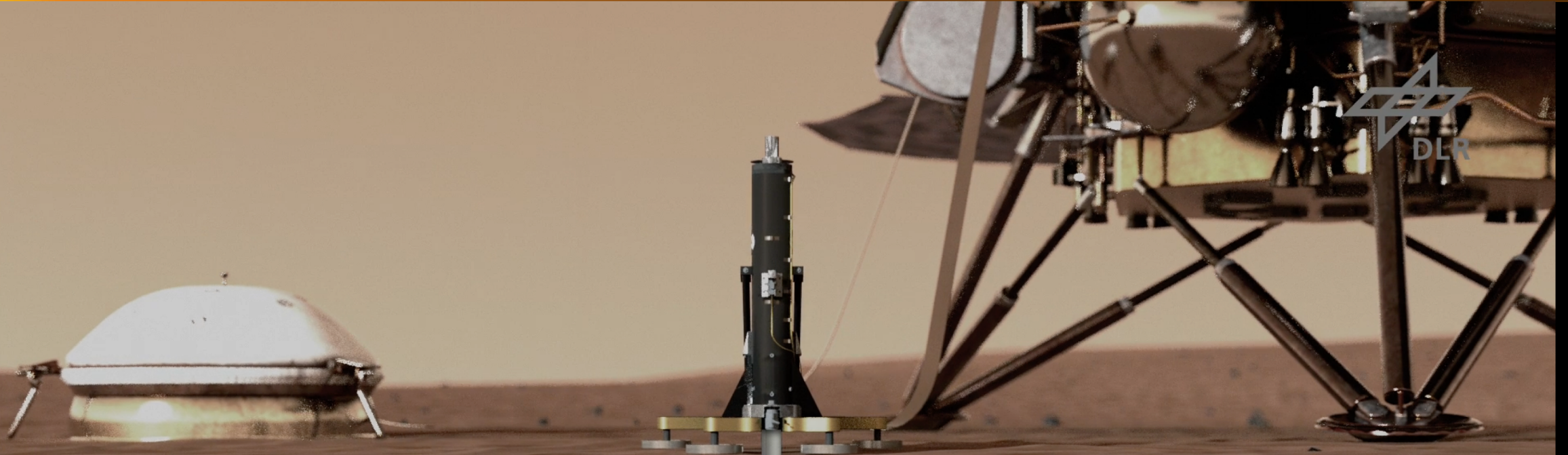


Instrument Electronics – Inside S/C
Pressure Sensor – Inside S/C
Radiometer – Other side of S/C
Camera Calibration Target – Other side of deck
LaRRI (Laser Retroreflector) – Other side of deck
Names to Mars Chips – Other side of deck

'Mole' penetrator

- ▮ Tilt meter
- ▮ Heated foils in the outer hull to measure the soil's thermal conductivity
- ▮ Hammering mechanism

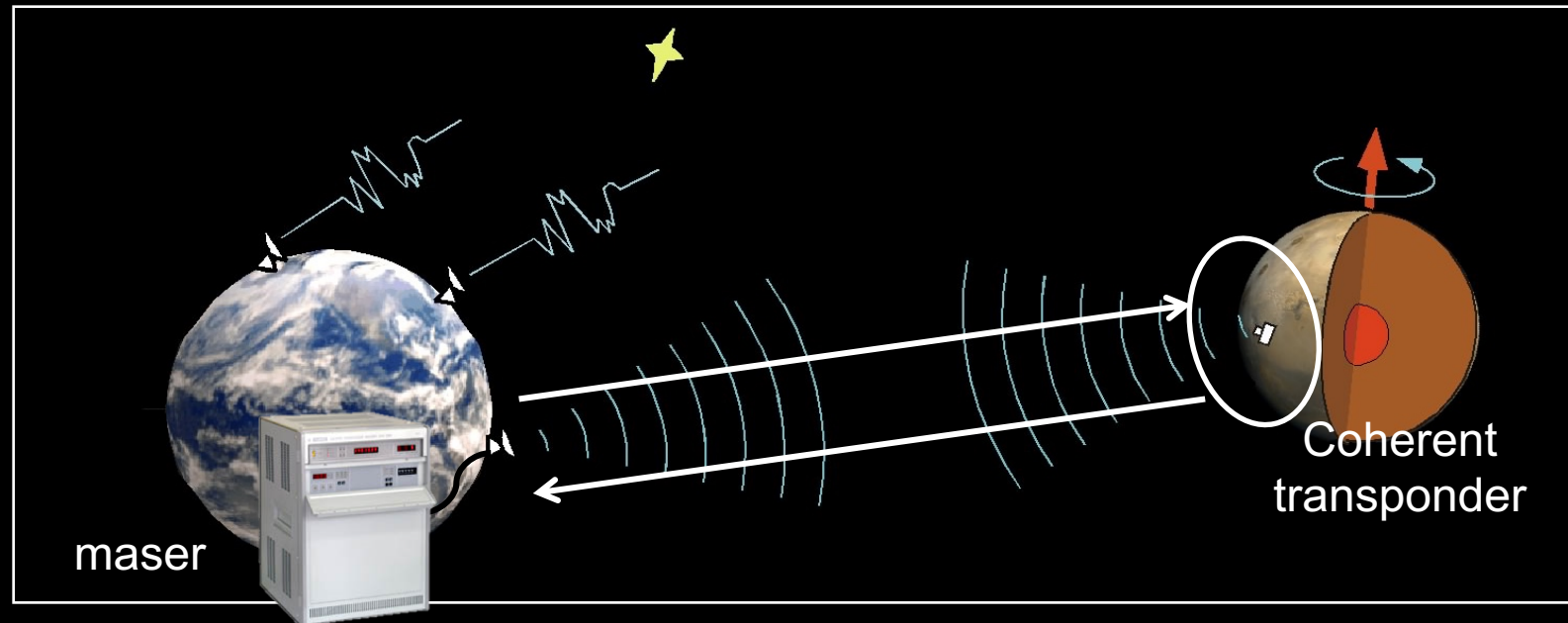


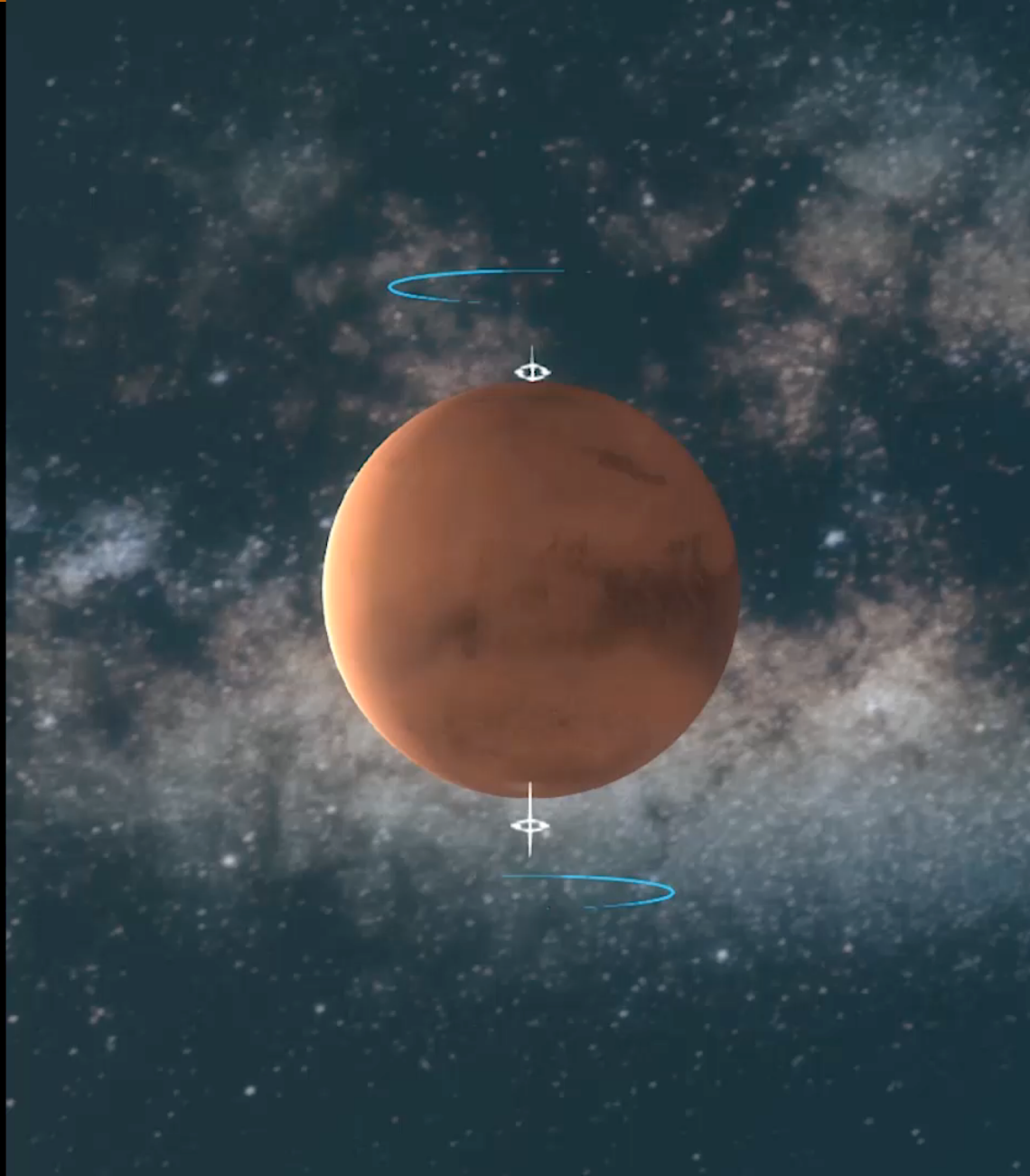


Hammering mechanism

- ▮ The primary spring is tensioned by a motor driven cam and roller. When the spring is released, it drives the hammer forward
- ▮ The recoil is absorbed by a second spring, and the friction on the hull

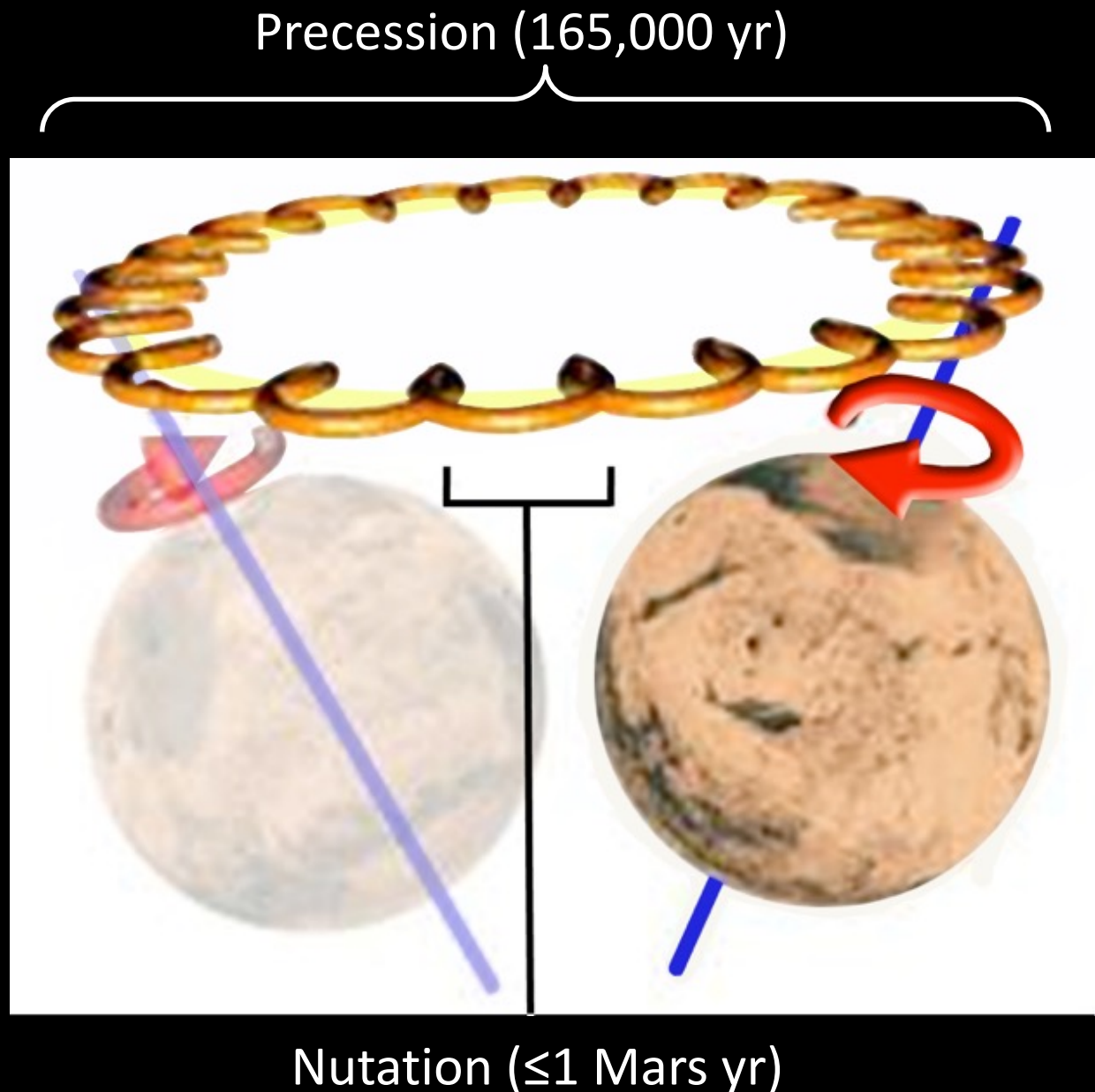
- Measurement of the timing and Doppler shift of the X-band radio signal between the Earth and InSight allow us to track the location and motion of the lander to an accuracy of better than 10 cm in inertial space.
- By tracking the lander location for about an hour each day, we will be able to determine the direction and motion of the rotation vector of Mars.







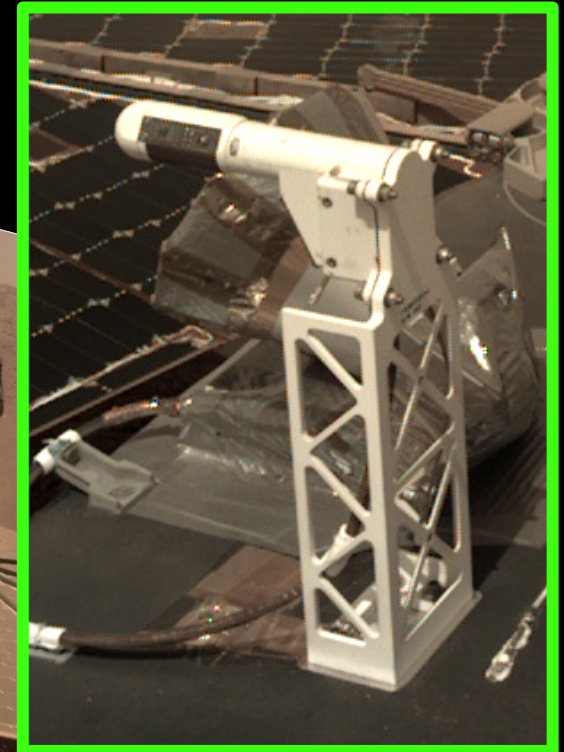
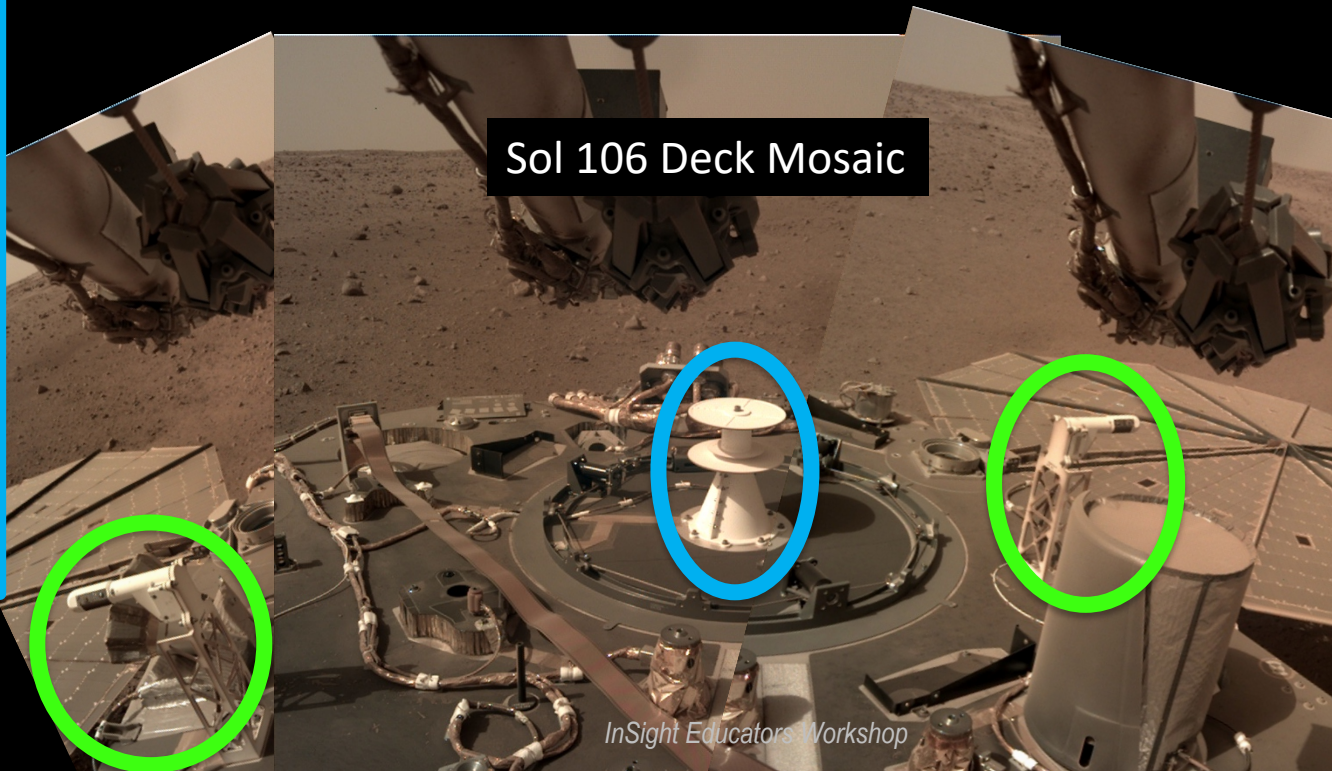
These two measurements taken together allow us to calculate the size and density of the metallic core of Mars.



Pressure Sensor



TWINS (Temperature & Wind for InSight)



Latest Weather at Elysium Planitia

InSight is taking daily weather measurements (temperature, wind, pressure) on the surface of Mars at Elysium Planitia, a flat, smooth plain near Mars' equator.

Sol 448
February 29

High: -8° F | C
Low: -94° F | C

Sol 442
Feb. 23

High: -10° C
Low: -94° C

Sol 443
Feb. 24

High: -12° C
Low: -95° C

Sol 444
Feb. 25

High: -12° C
Low: -94° C

Sol 445
Feb. 26

High: -13° C
Low: -93° C

Sol 446
Feb. 27

High: -11° C
Low: -94° C

Sol 447
Feb. 28

High: -11° C
Low: -93° C

Sol 448
Feb. 29

High: -8° C
Low: -94° C

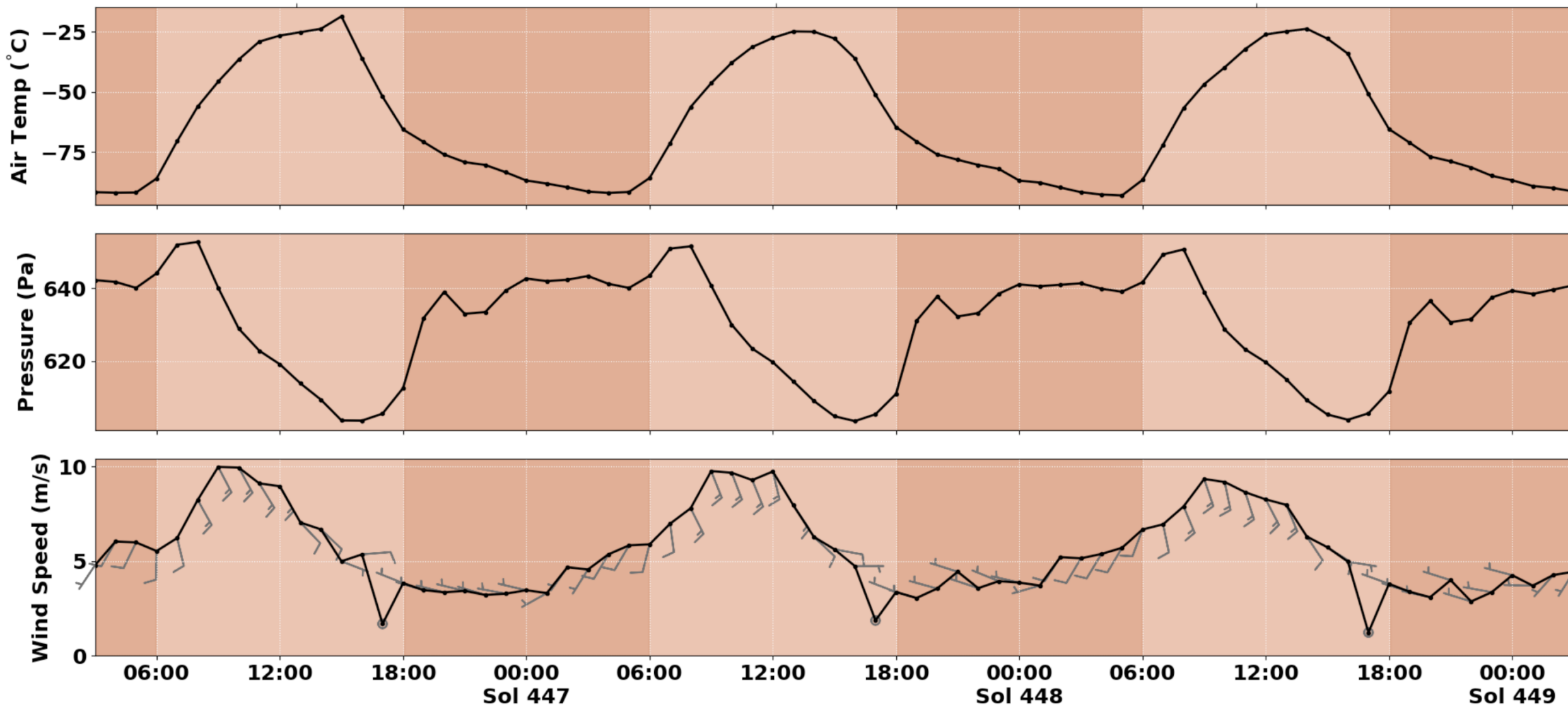


Pressure, Temperature and Wind Plots

Earth UTC:
02/28 00:00

02/29 00:00

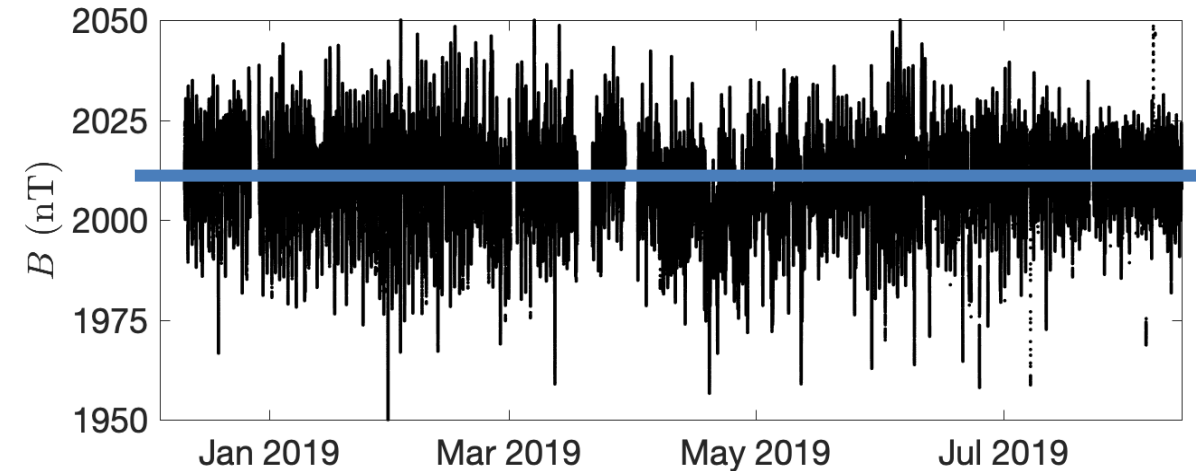
03/01 00:00



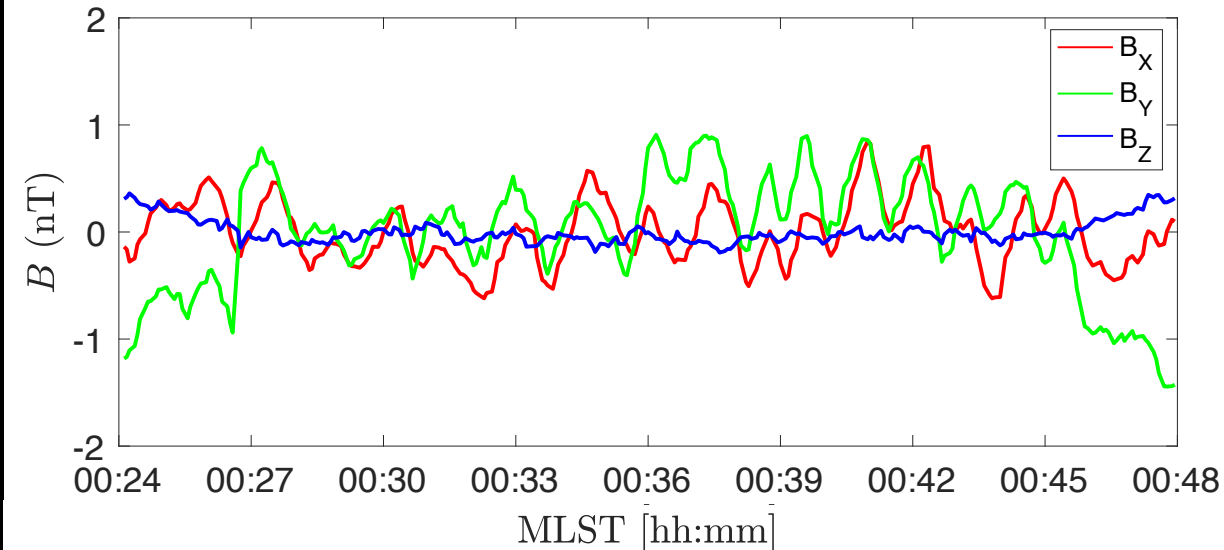
First Magnetic Measurements from the Surface of Mars

- The constant magnetic field at the landing site is about 10X stronger than measured from orbit
 \Rightarrow significant crustal variations at spatial scales < 150 km.
- Midnight pulsations are observed that are probably due to electric currents in the ionosphere of Mars.
 \Rightarrow may be used to probe the conductivity of Mars as a function of depth.

All IFG Data Through July 2019

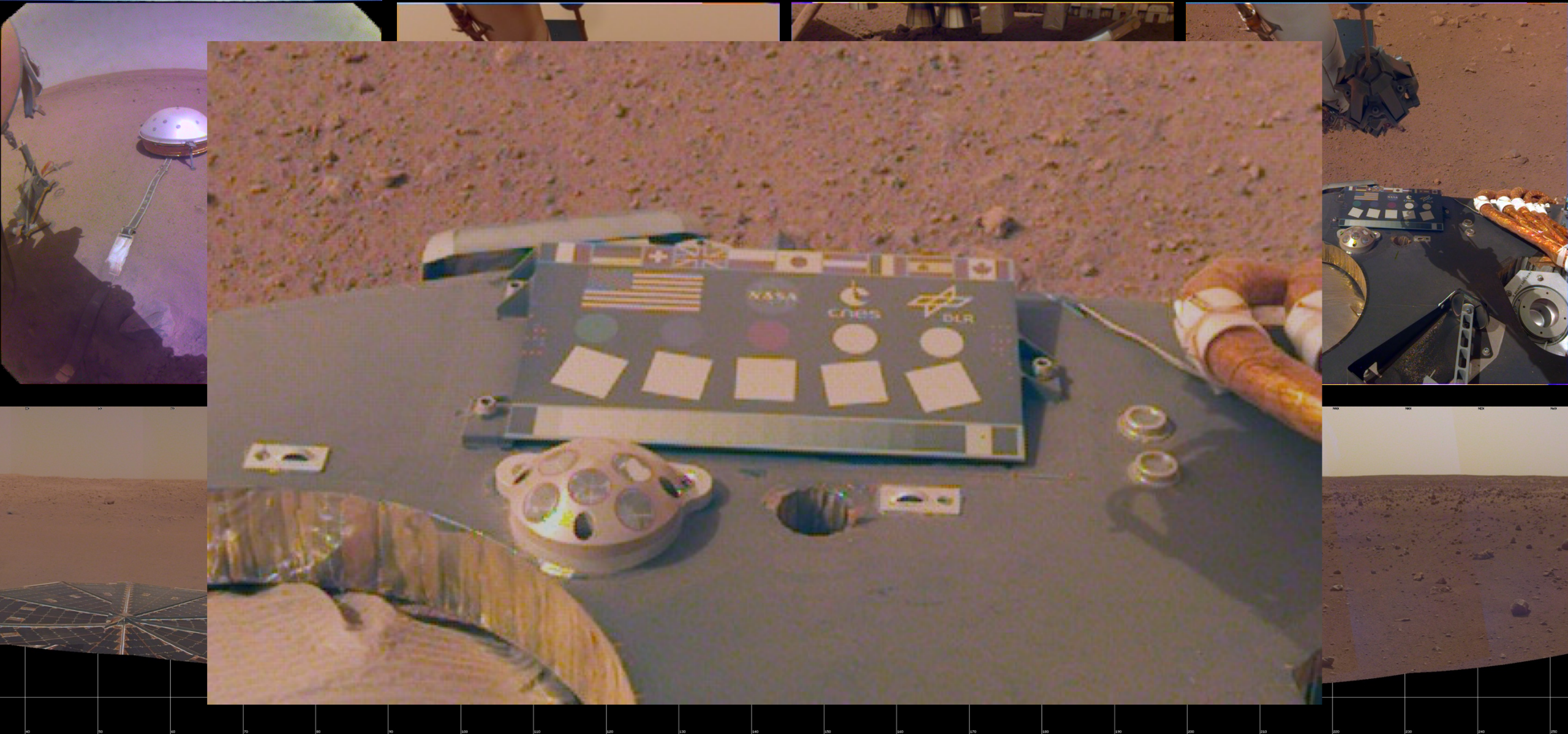


SOL16





Pictures! (4127 Images as of Sol 448, 1 March, 2020)



Thanks for listening!

See all the latest pictures and weather
(posted as soon as the data hits the
ground!) plus other InSight news at

mars.nasa.gov/insight